A novel search for gravitational waves inspired by axion dark matter

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Novel Search for High-Frequency Gravitational Waves with Low-Mass Axion Haloscopes

Valerie Domcke, Camilo Garcia-Cely, and Nicholas L. Rodd Phys. Rev. Lett. **129**, 041101 – Published 20 July 2022

Outline

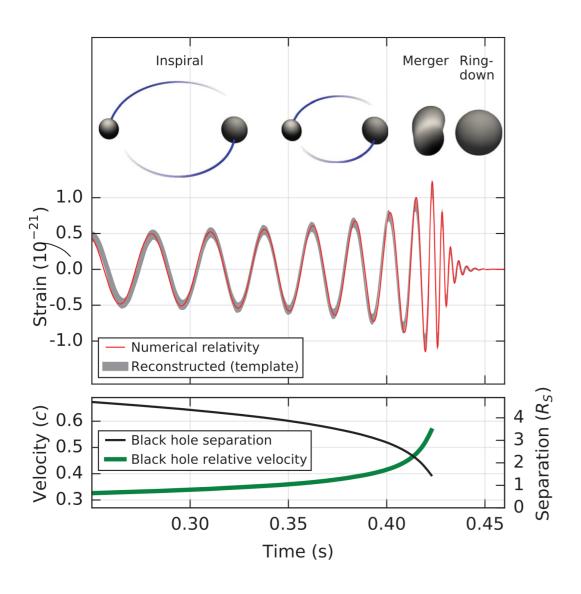
Motivation

Adapting axion experiments to detect gravitational waves

Conclusions

Motivation

Gravitational waves from black holes



PRL 116, 061102 (2016)

PHYSICAL REVIEW LETTERS

12 FEBRUARY 2016

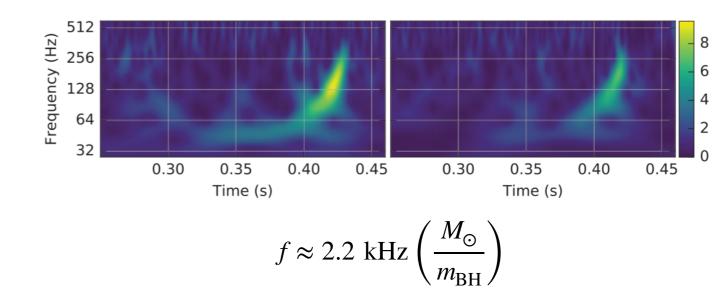


Observation of Gravitational Waves from a Binary Black Hole Merger

B. P. Abbott *et al.**

(LIGO Scientific Collaboration and Virgo Collaboration)

(Received 21 January 2016; published 11 February 2016)



Subsolar primordial black holes can produce high frequency GWs

High-frequency gravitational waves

No known astrophysical objects are small and dense enough to produce gravitational waves beyond 10 kHz

Part of a collection:

Gravitational Waves

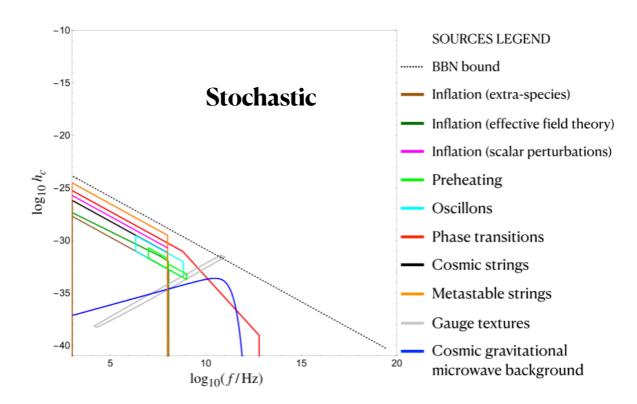
Review Article | Open Access | Published: 06 December 2021

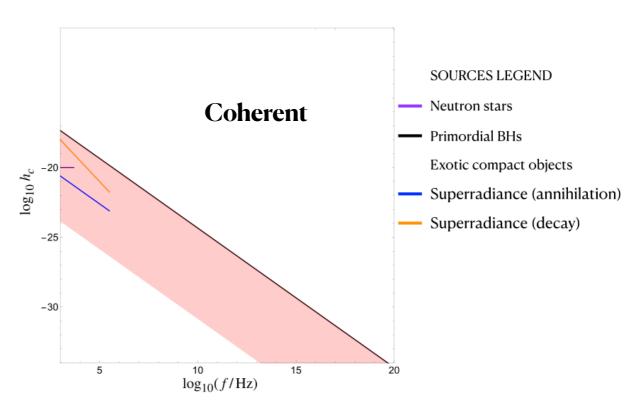
Challenges and opportunities of gravitational-wave searches at MHz to GHz frequencies

Nancy Aggarwal , Odylio D. Aguiar, Andreas Bauswein, Giancarlo Cella, Sebastian Clesse, Adrian Michael Cruise, Valerie Domcke, Daniel G. Figueroa, Andrew Geraci, Maxim Goryachev, Hartmut Grote, Mark Hindmarsh, Francesco Muia, Nikhil Mukund, David Ottaway, Marco Peloso, Fernando Quevedo, Angelo Ricciardone, Jessica Steinlechner, Sebastian Steinlechner, Sichun Sun, Michael E. Tobar, Francisco Torrenti, Caner Ünal & Graham White

Living Reviews in Relativity 24, Article number: 4 (2021) | Cite this article

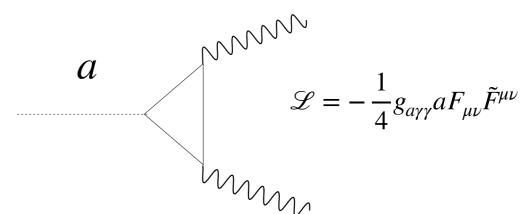
A growing community is seriously considering the search of high frequency gravitational waves





Axion dark matter

Pseudoscalar field



Solution to the strong CP problem

Peccei, Quinn 1977

• Excellent dark matter candidate

Weinberg, Wilczek 1978

Effective current for axions

Axions act as a source term to Maxwell's equations, effectively inducing an electromagnetic current.

$$\nabla \cdot \mathbf{B} = 0$$

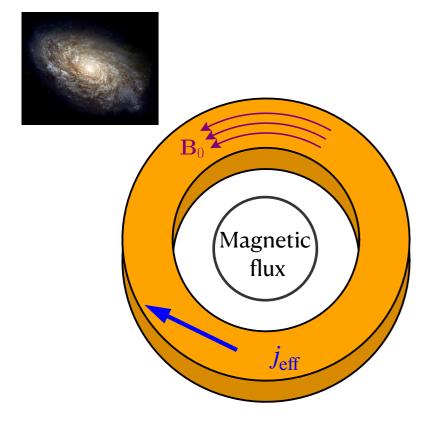
$$\nabla \times \mathbf{E} + \partial_t \mathbf{B} = 0$$

$$\nabla \cdot \mathbf{E} = j^0$$

$$\nabla \times \mathbf{B} - \partial_t \mathbf{E} = \mathbf{j}$$

$$j^{0} = -g_{a\gamma\gamma} \nabla a \cdot \mathbf{B}$$
 $\mathbf{j} = g_{a\gamma\gamma} (\nabla a \times \mathbf{E} + \partial_{t} a \mathbf{B})$

Low mass axion haloscopes

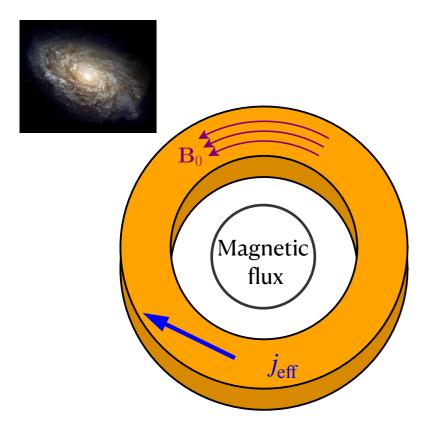


$$\nabla \times \mathbf{B} - \partial_t \mathbf{E} = g_{a\gamma\gamma} \, \partial_t a \, \mathbf{B_0}$$

$$j_{\text{eff}}$$

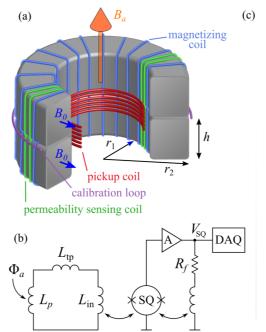
The electromagnetic fields produced by the axion drive a current through a pickup coil

Low mass axion haloscopes



$$\nabla \times \mathbf{B} - \partial_t \mathbf{E} = g_{a\gamma\gamma} \, \partial_t a \, \mathbf{B_0}$$

$$j_{\text{eff}}$$





Search for axion-like dark matter with ferromagnets

Alexander V. Gramolin^{⊙1}, Deniz Aybas^{⊙1,2}, Dorian Johnson¹, Janos Adam¹ and Alexander O. Sushkov^{⊙1,2,3} ⊠



PRL 117, 141801 (2016)

PHYSICAL REVIEW LETTERS

week ending 30 SEPTEMBER 201

Broadband and Resonant Approaches to Axion Dark Matter Detection

Yonatan Kahn, ^{1,*} Benjamin R. Safdi, ^{2,†} and Jesse Thaler ^{2,‡}

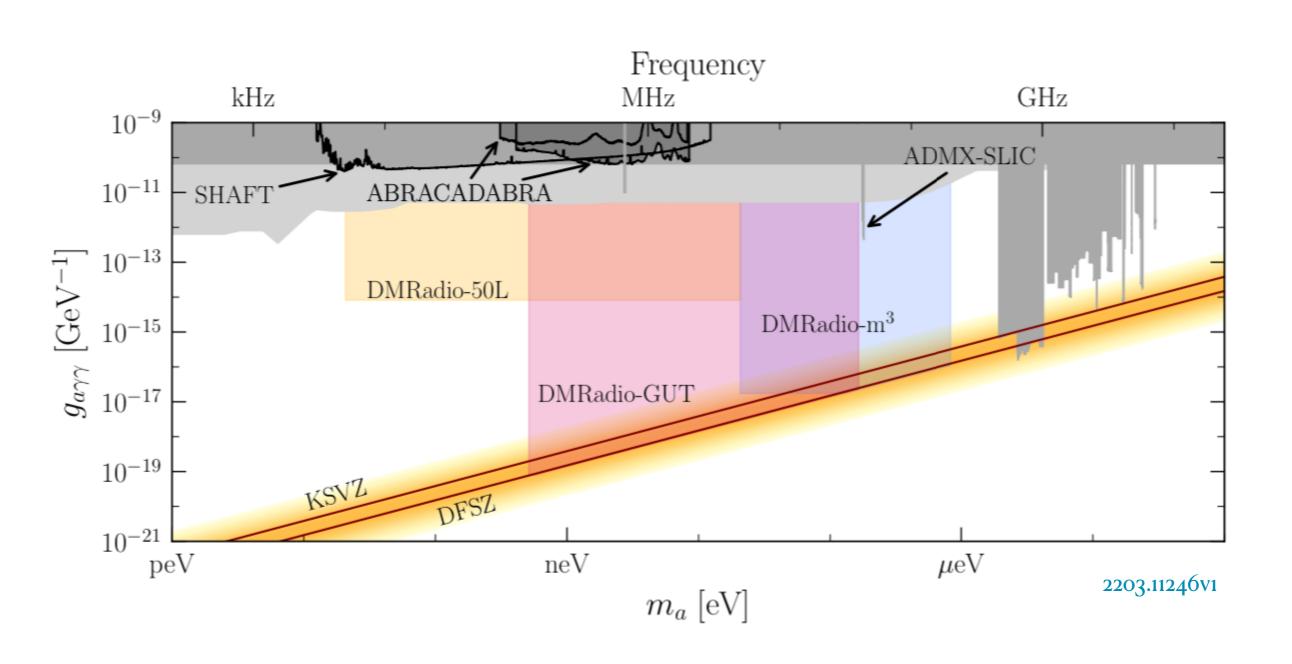
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²Center for Theoretical Physics, Massachusetts Institute of Technology, Cambridge, Massachusetts 02139, USA

(Received 3 March 2016; published 30 September 2016)

The electromagnetic fields produced by the axion drive a current through a pickup coil

DMRadio program

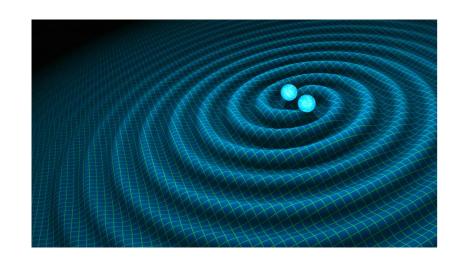


Adapting axion experiments to detect gravitational waves

Effective current for gravitational waves

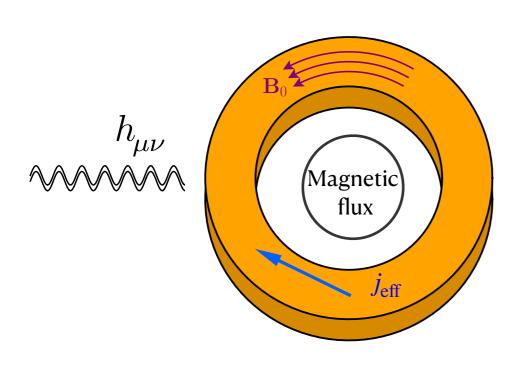
GWs act as a source term to Maxwell's equations, effectively inducing an electromagnetic current.

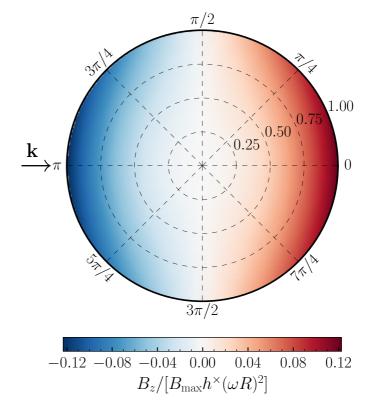
$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu} \quad \left| h_{\mu\nu} \right| \ll 1$$



$$j_{\text{eff}}^{\mu} = \partial_{\nu} \left(-\frac{1}{2} h F^{\mu\nu} + F^{\mu\alpha} h^{\nu}_{\alpha} - F^{\nu\alpha} h^{\mu}_{\alpha} \right)$$

Domcke, CGC, Rodd, 2202.00695



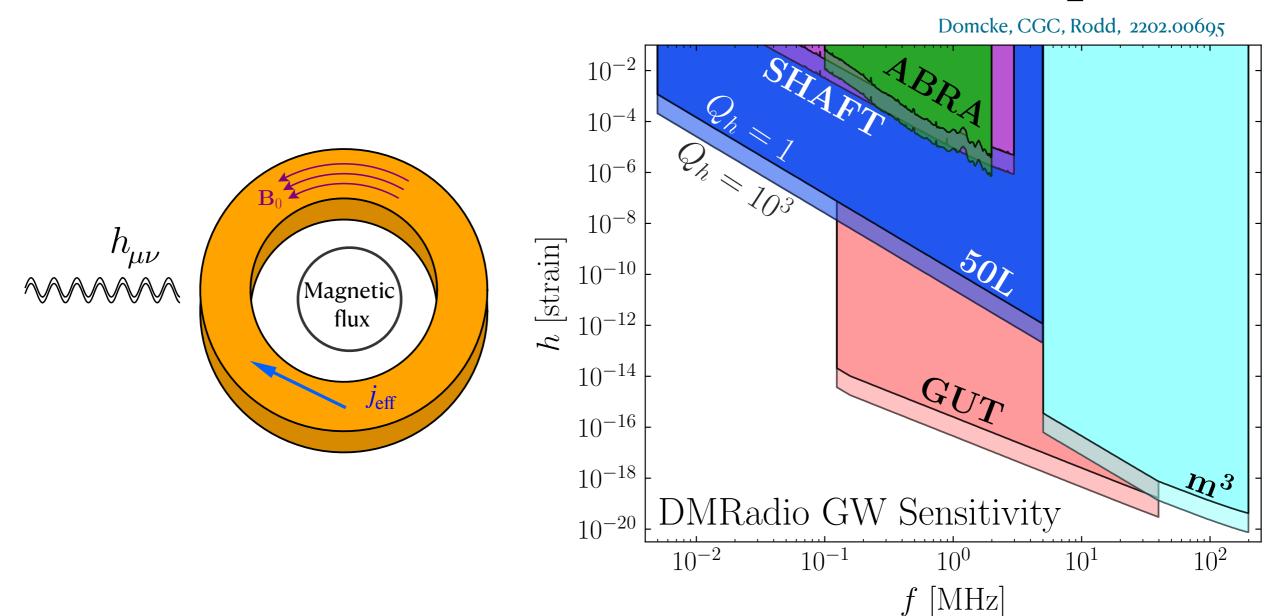


$$\Phi \approx \frac{\mathrm{i}e^{-\mathrm{i}\omega t}}{16\sqrt{2}}h^{\times}\omega^{3}B_{\mathrm{max}}\pi r^{2}Ra(a+2R)s_{\theta_{h}}^{2}$$

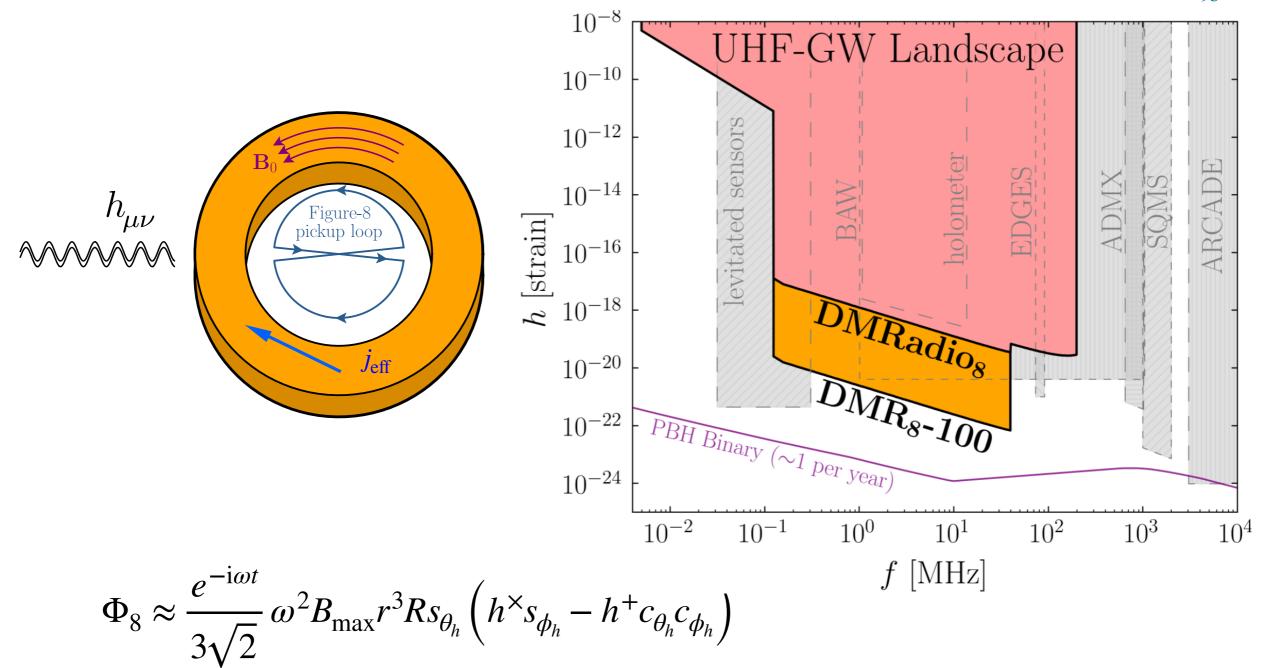
Suppression at small frequencies

$$\Phi_{\rm axions} \approx e^{-i\omega t} g_{a\gamma\gamma} \sqrt{2\rho_{\rm DM}} B_{\rm max} \pi r^2 R$$

The sensitivity scaling with the volume is faster than for axions

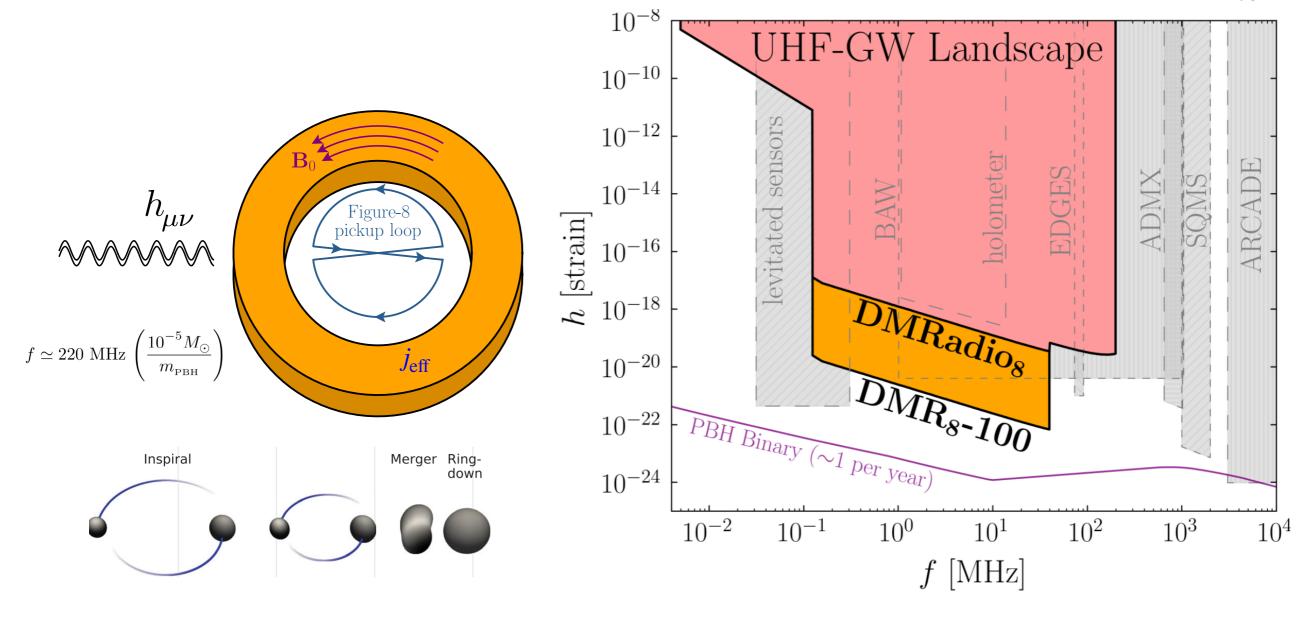


Domcke, CGC, Rodd, 2202.00695



Small modification allows to measure both polarizations

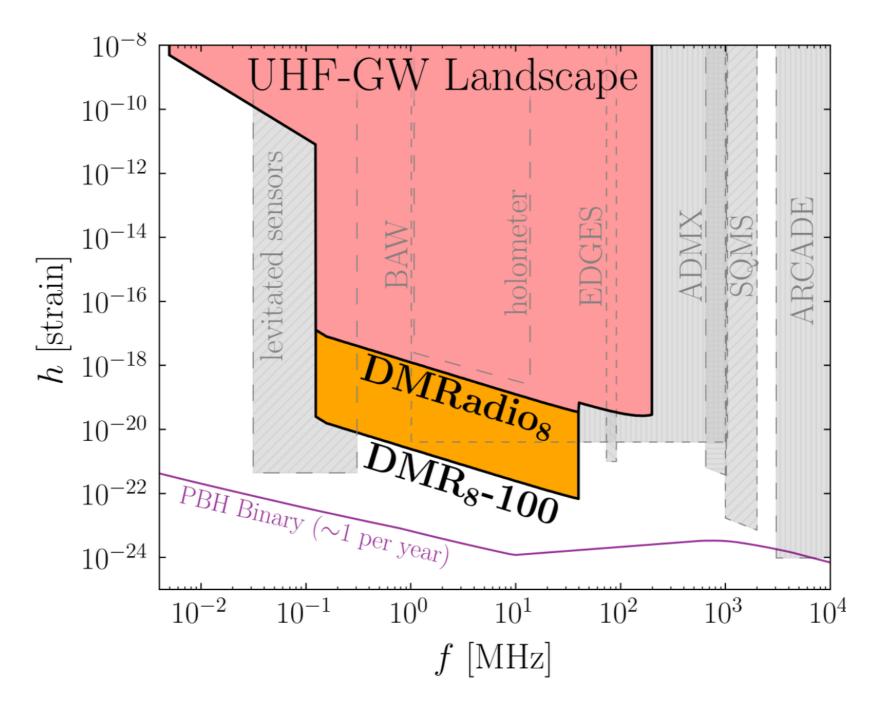
Domcke, CGC, Rodd, 2202.00695



Up-to-date estimate of PBH in binaries and their expected merger rate accounting for the local overdensity in the Milky Way

See also 2205.02153 by Franciolini, A. Maharana, and F. Muia,

Conclusions



Axion experiments may discover not only dark matter, but also exotic sources of gravitational waves

Different experimental proposals have coalesced on a strain sensitivity of 10^{-22} for MHz GWs, still orders of magnitude away from signals of the early Universe. Whether we can hope to probe such strain sensitivities remains to be determined.

The nucleosynthesis bound

A. Ejlli ☑, D. Ejlli, A. M. Cruise, G. Pisano & H. Grote

The European Physical Journal C 79, Article number: 1032 (2019)

