

Computer tools in particle physics

- Introduction -

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Back in the good old times...

Dear radioactive Ladies and Gentlemen...

Physikalisches Institut
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Zürich

Zürich, 4. Dez. 1930
Gloriastrasse

Liebe Radioaktive Damen und Herren,

Wie der Ueberbringer dieser Zeilen, den ich halbvollst anzuhören bitte, Ihnen des näheren auseinandersetzen wird, bin ich angesichts der "falschen" Statistik der N- und Li-6 Kerne, sowie des kontinuierlichen beta-Spektrums auf einen verweifelten Ausweg verfallen um den "Wechselsatz" (1) der Statistik und den Energiesatz zu retten. Nämlich die Möglichkeit, es könnten elektrisch neutrale Teilchen, die ich Neutronen nennen will, in den Kernen existieren, welche den Spin $1/2$ haben und das Ausschliessungsprinzip befolgen und sich von Lichtquanten ausserdem noch dadurch unterscheiden, dass sie nicht mit Lichtgeschwindigkeit laufen. Die Masse der Neutronen müsste von derselben Grössenordnung wie die Elektronenmasse sein und jedenfalls nicht grösser als $0,01$ Protonenmasse.- Das kontinuierliche beta-Spektrum wäre dann verständlich unter der Annahme, dass beim beta-Zerfall mit dem Elektron jeweils noch ein Neutron emittiert wird, derart, dass die Summe der Energien von Neutron und Elektron konstant ist.



December 4th, 1930

Letter to his colleagues in Tübingen

1930

Pauli's neutrino hypothesis

Back in the good old times...

Zürich, Dec. 4, 1930

Physics Institute of the ETH

Gloriastrasse

Zürich

Dear Radioactive Ladies and Gentlemen,

As the bearer of these lines, to whom I graciously ask you to listen, will explain to you in more detail, because of the "wrong" statistics of the N- and Li-6 nuclei and the continuous beta spectrum, I have hit upon a desperate remedy to save the "exchange theorem" of statistics and the law of conservation of energy. Namely, the possibility that in the nuclei there could exist electrically neutral particles, which I will call neutrons, that have spin $1/2$ and obey the exclusion principle and that further differ from light quanta in that they do not travel with the velocity of light.

(.../...)

But so far I do not dare to publish anything about this idea, and trustfully turn first to you, dear radioactive people, with the question of how likely it is to find experimental evidence for such a neutron if it would have the same or perhaps a 10 times larger ability to get through [material] than a gamma-ray.

I admit that my remedy may seem almost improbable because one probably would have seen those neutrons, if they exist, for a long time. (.../...) Thus, dear radioactive people, scrutinize and judge. - Unfortunately, I cannot personally appear in Tübingen since I am indispensable here in Zürich because of a ball on the night from December 6 to 7. With my best regards to you, and also to Mr. Back, your humble servant

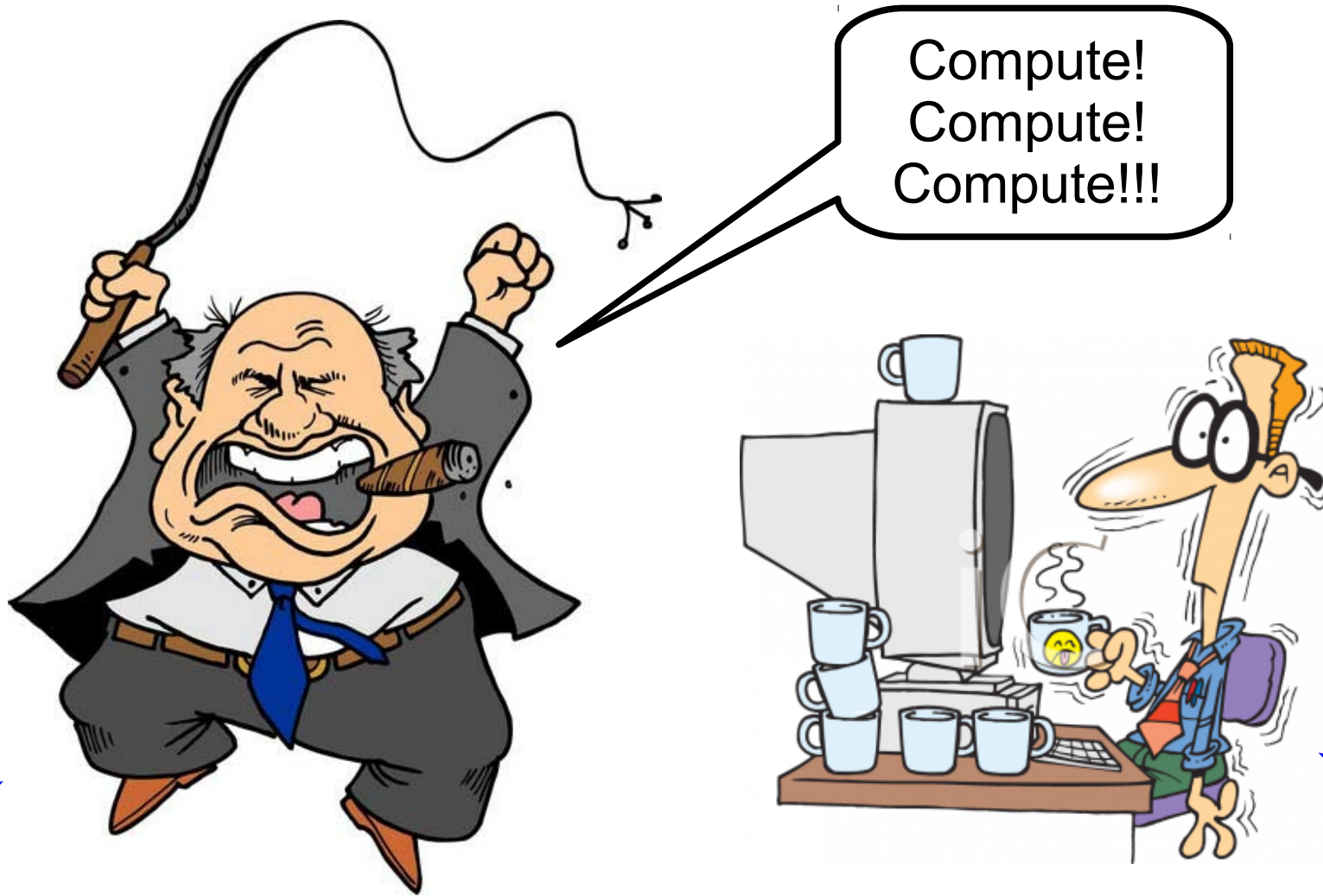
signed W. Pauli

Many new models and particles

An “explosion” of new **models** and **particles**. Strategy:

- **Analytical** derivation of particle masses and vertices, minimization of the scalar potential, renormalization group equations, ...
- **Numerical** routines: diagonalization, resolution of differential equations, phase space integration...
- Mass spectrum, loop corrections, flavor **observables** and decay rates
- **Dark matter** properties: relic density, direct and indirect detection rates, ...
- **Collider** simulations
- **Other**

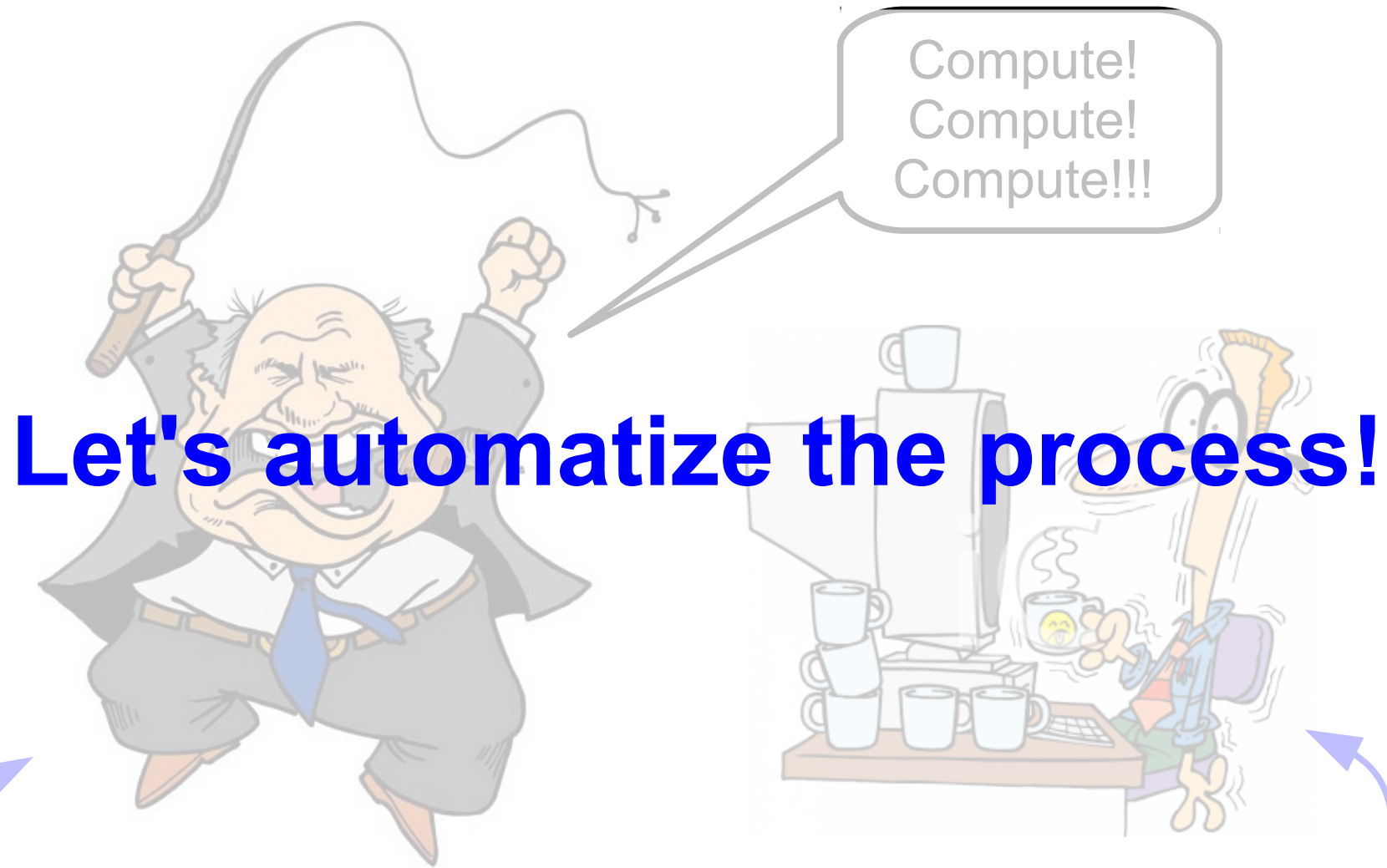
Usual approach



Professor

Poor student

Usual approach



Professor

Poor student

SARAH and SPheno

SARAH



[Staub]

SARAH is a **Mathematica package** for analyzing SUSY and non-SUSY models.

It calculates **analytically** all vertices, mass matrices, tadpoles equations, 1-loop corrections for tadpoles and self-energies and 2-loop RGEs.

SARAH is also a spectrum-generator-generator: based on the derived analytical expressions it creates Fortran source code for **SPheno**.

SPheno

[Porod, Staub]

SPheno is a **Fortran code**. It provides routines for the **numerical evaluation** of all vertices, masses and decay modes in a given model.

MicrOmegas and MadGraph

MicrOmegas

[Bélanger, Boudjema, Pukhov, Semenov]

Computer code for the study of **dark matter**.

First developed to compute the **relic density** of a stable massive particle, the current version also computes direct and indirect dark matter detection rates.

Written in **C** and **Fortran**.

MadGraph

[The MadTeam]

MadGraph is a Monte Carlo event generator for collider simulations. It allows for a complete simulation of a new physics model at the **LHC**, from events at the parton level to detector response.

It is written in **Python**.

Message 1

It is not so hard!



**What people think about
SARAH, micrOmegas,
MadGraph...**



What they really are

Message 2

Do no trust (too much) in codes!



Plan

- **Lecture 1** : Exploring new models with SARAH
- **Lecture 2** : Computing dark matter properties with MicrOmegas
- **Lecture 3** : LHC physics with MadGraph
- **Lecture 4** : Final exercise
- **Lecture 5** : Introduction to other codes, questions and conclusion



References

Lectures

“Computer tools in particle physics”, A. Vicente,
[\[arXiv:1507.06349\]](#)

Practical introductions

“Exploring new models in all detail with SARAH”, F. Staub,
[\[arXiv:1503.04200\]](#)

(Only for SUSY)

Manuals

SARAH: [arXiv:1309.7223](#)

Spheno: [arXiv:1104.1573](#)

MicrOmegas: [arXiv:1407.6129](#)

MadGraph: [arXiv:1405.0301](#)

Websites

All links can be found in

<http://ific.uv.es/~montesin/ific2017.html>



Let's get started!



Rules:

- You can interrupt and ask questions at any moment
- Suggestion: you can emulate what I do with your own laptop
- I will assume that you already have all the prerequisites installed