

# Bounds on Neutrino-Scalar Coupling

Pedro S. Pasquini, Orlando L. G. Peres

Universidade Estadual de Campinas (Unicamp)

**Encontro Nacional de Física de Partículas e Campos**

14 a 18 Setembro 2015, Caxambu, MG

# SM is not the final chapter!

The **PARTICLE ZOO** Seeing the fabric of spacetime

ELEMENTARY PARTICLES of THE STANDARD MODEL:

	FERMIONS			BOSONS
	I	II	III	
QUARKS	 $u$ UP QUARK	 $c$ CHARM QUARK	 $t$ TOP QUARK	 $\gamma$ PHOTON
	 $d$ DOWN QUARK	 $s$ STRANGE QUARK	 $b$ BOTTOM QUARK	 $g$ GLUON
	 $\nu_e$ ELECTRON-NEUTRINO	 $\nu_\mu$ MUON-NEUTRINO	 $\nu_\tau$ TAU-NEUTRINO	 $Z$ Z BOSON
LEPTONS	 $e^-$ ELECTRON	 $\mu$ MUON	 $\tau$ TAU	 $W$ W BOSON

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Neutrinos have Mass!

And we don't know why

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Perfect to look for new interactions/particles!

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## $\nu$ -Scalar interactions are interesting!

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$$\mathcal{L}_{\text{Yukawa}} = \frac{1}{2}g_{\alpha\beta}\bar{\nu}_\alpha\nu_\beta\phi_1 + \frac{i}{2}g'_{\alpha\beta}\bar{\nu}_\alpha\gamma_5\nu_\beta\phi_2$$

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This interactions can be found in many models, e. g.

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Two Higgs Models [2]:

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Meson 

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lepton ←

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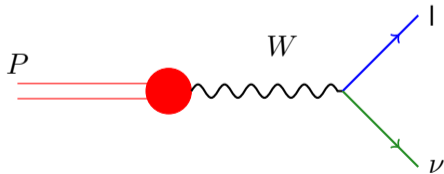
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Neutrino

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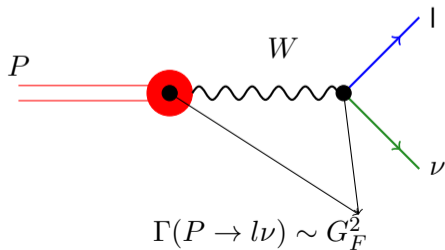
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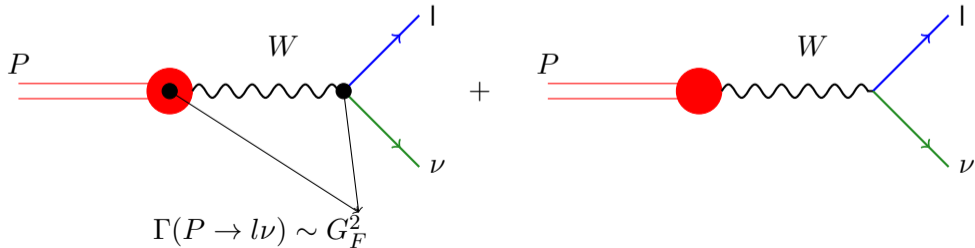




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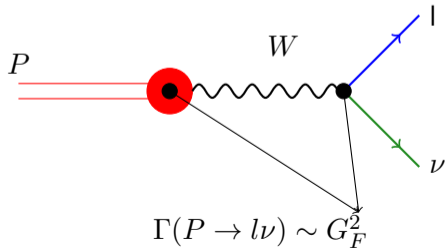
$$P \rightarrow l + \nu_m + \phi$$



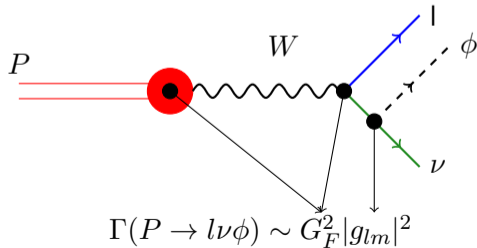
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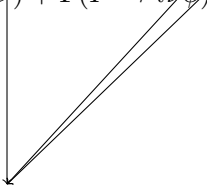
# New couplings are small Corrections to known physics

$$\Gamma_{\text{tot}} = \Gamma(P \rightarrow l\nu) + \Gamma(P \rightarrow l\nu\phi)$$

Decay Rate ( $P$  rest Frame)

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Can't be observed on this experiment

## New couplings are small Corrections to known physics

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SM  
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New Physics

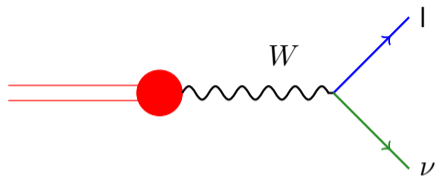
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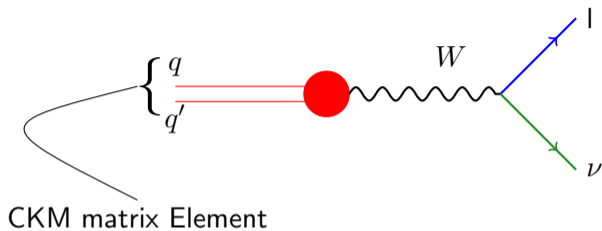
$$\Gamma(P \rightarrow l\nu) = \frac{G_F^2 f_P^2 |V_{qq'}|^2 m_p^3}{8\pi} \left(\frac{m_l}{m_p}\right)^2 \left[1 - \left(\frac{m_l}{m_p}\right)^2\right]^2$$

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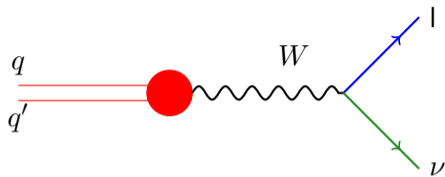
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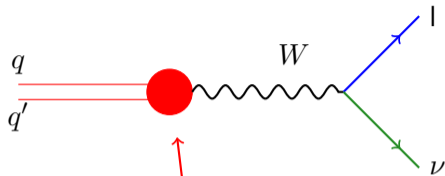
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Chiral suppression

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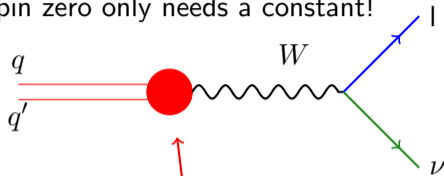


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Meson Decay Constant

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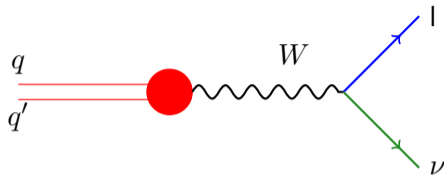
Spin zero only needs a constant!



$$\Gamma(P \rightarrow l\nu) = \frac{G_F^2 f_P^2 |V_{qq'}|^2 m_P^3}{8\pi} \left(\frac{m_l}{m_p}\right)^2 \left[1 - \left(\frac{m_l}{m_p}\right)^2\right]^2$$

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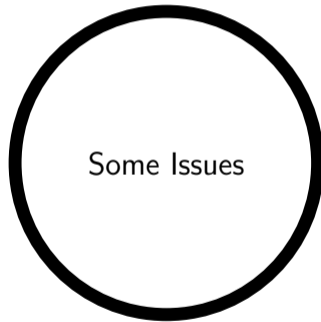


More on that later...



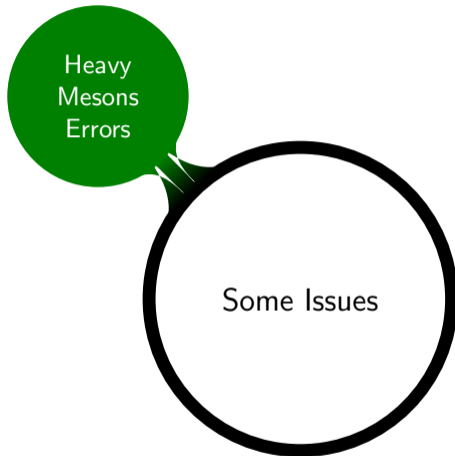
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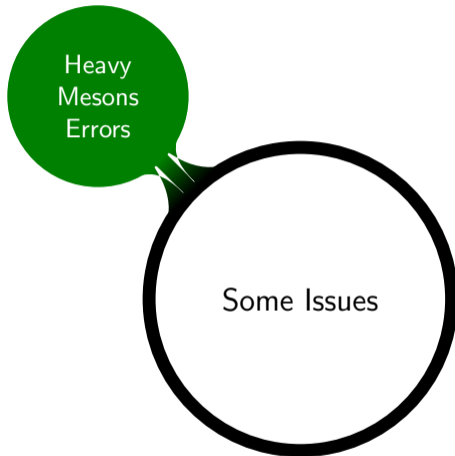


# We don't know everything!

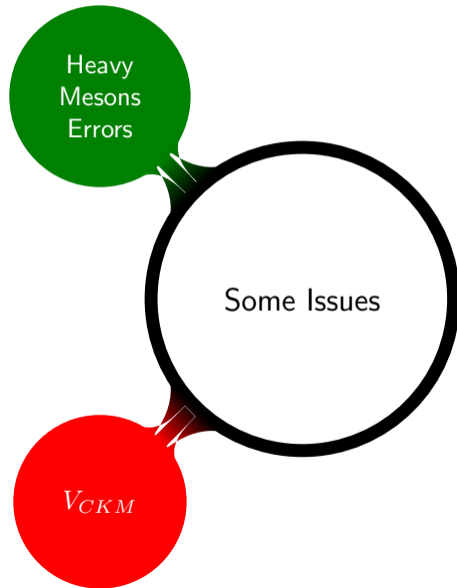


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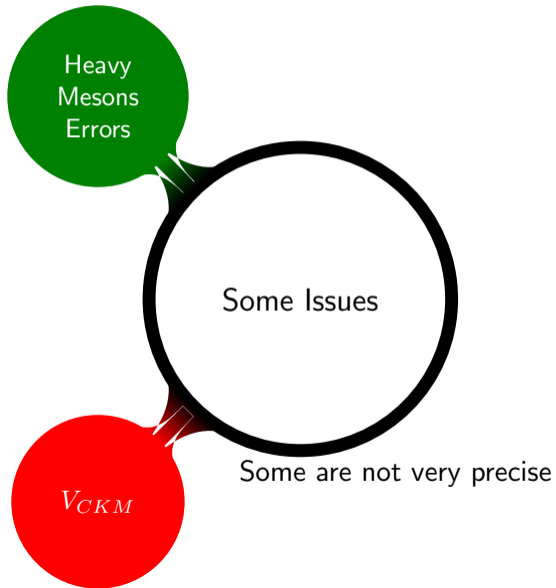
Reach up to 25%...



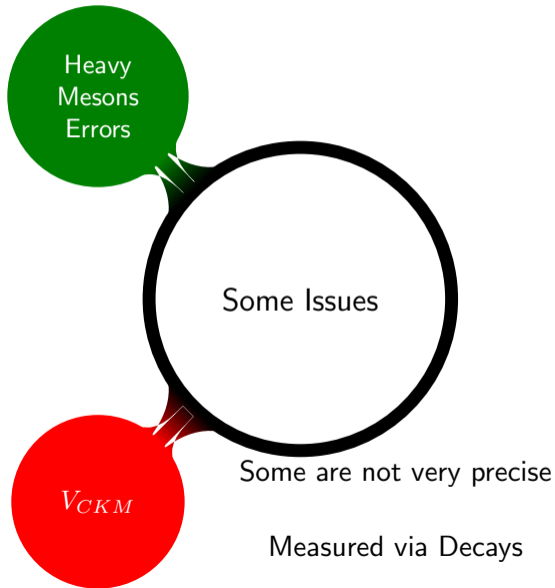
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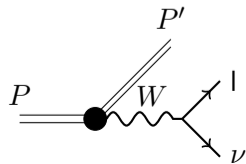
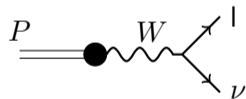
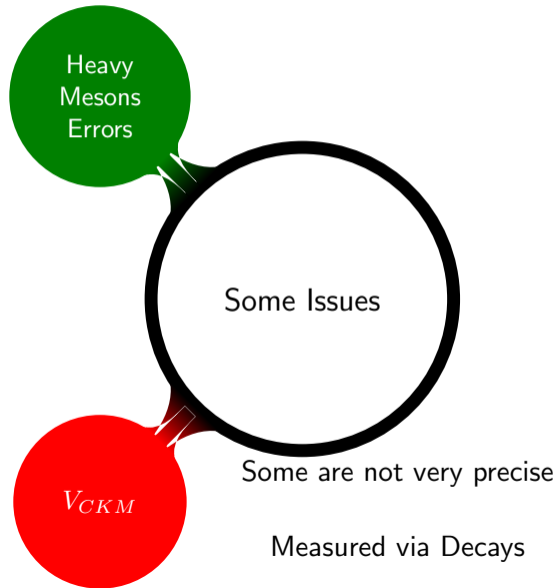
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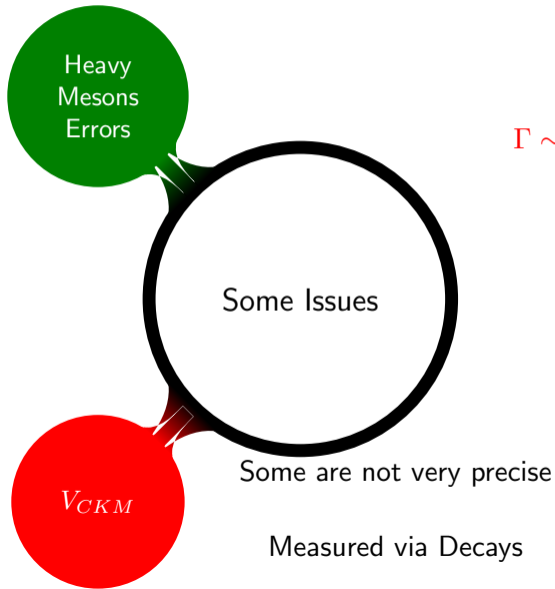
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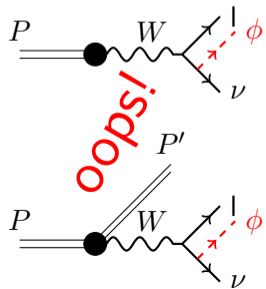
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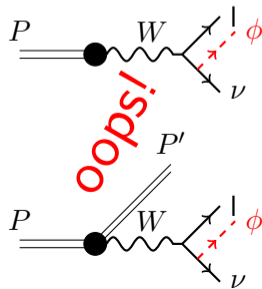
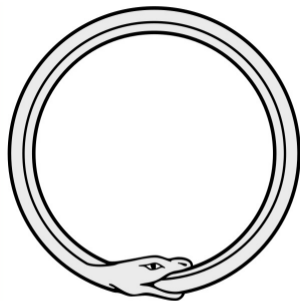
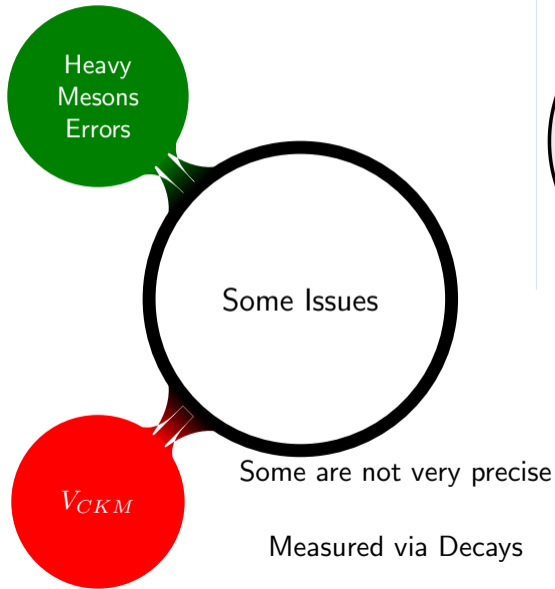
# We don't know everything!



$$\Gamma \sim |V_{CKM}|^2(1 + \text{corrections})$$

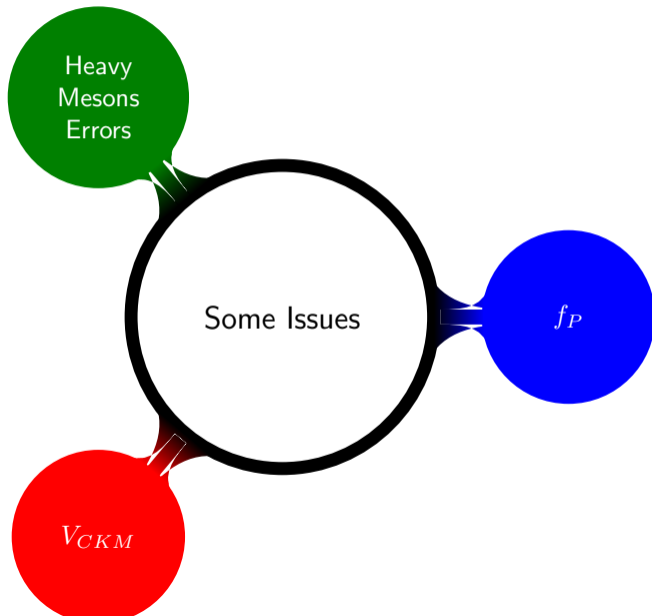


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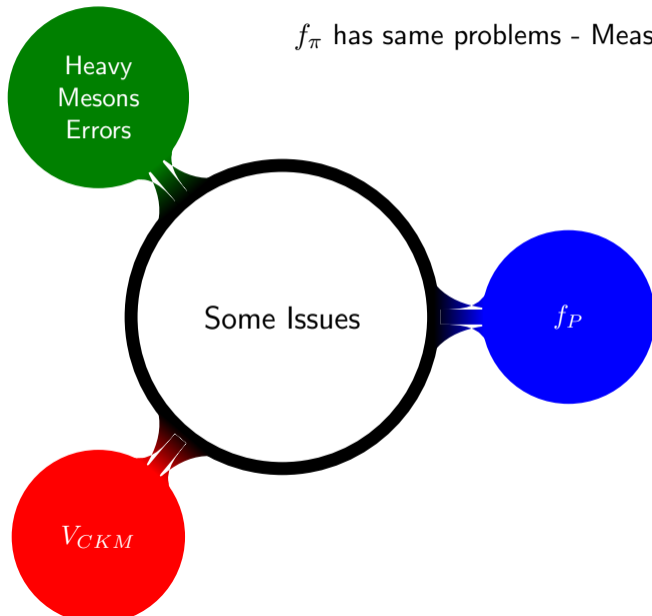


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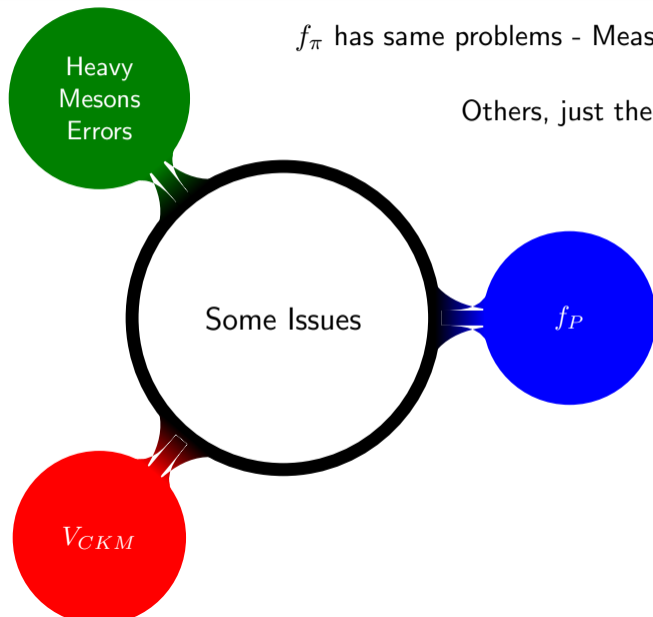


# We don't know everything!

$f_\pi$  has same problems - Measured via decays



# We don't know everything!



$f_\pi$  has same problems - Measured via decays

Others, just theoretical!

Lattice QCD<sup>3</sup>

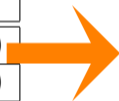
<sup>3</sup> Eur. Phys. J. C **74**, no. 9, 2890 (2014)

## Lattice QCD<sup>3</sup>

	$f_P[\text{MeV}]$
$\pi$	130.2(1.4)
$K$	156.3(0.9)
$D$	209(3.3)
$D_s$	250(7)
$B$	186(4)

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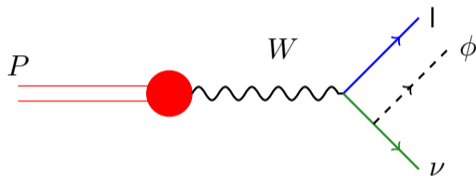
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Very Recent!

# It is a Three body Decay

Finally, Beyond SM:



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Following Barger, Keung and Pakvasa<sup>4</sup>



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Calculated the corrections  $A_{Pl}(m_\phi)$

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$$d\Gamma(P \rightarrow l\nu\phi) = \Gamma_{l\nu} \frac{(x^2 + \epsilon^2 + 6x\epsilon - \beta x - \beta\epsilon)\lambda^{1/2}(x, \epsilon, \beta)}{x^2(x-\epsilon)^2} \frac{g_l^2}{32\pi^2} dx$$


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$$\lambda(x, y, z) = x^2 + y^2 + z^2 - 2xy - 2xz - 2zy$$

Kinematic Function

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Leptonic Decay rate ←

for neutrino with Squared mass  $x$

$$\Gamma_{l\nu} = \frac{G_F^2 f_P^2 |V_{qq'}|^2 m_P^3}{8\pi} [x + \alpha - (x - \alpha)^2] \lambda^{1/2}(1, x, \alpha)$$

$$\alpha = \frac{m_l^2}{m_P^2}$$

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Three body and Yukawa

<sup>4</sup>Phys. Rev. D **25**, 907 (1982).

It is like a virtual twobody!

$x$  varies angles and  $M_{\text{miss}}^2$

# It is Divergent!

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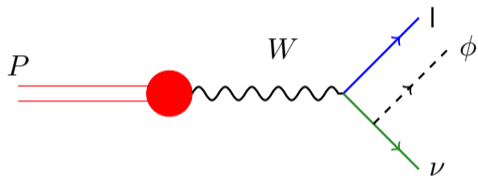
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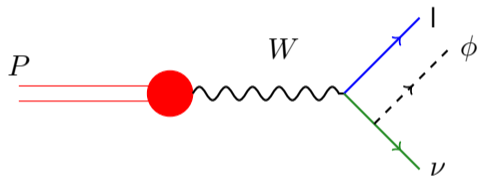
Infrared Divergent!!

<sup>4</sup>Phys. Rev. D **25**, 907 (1982).

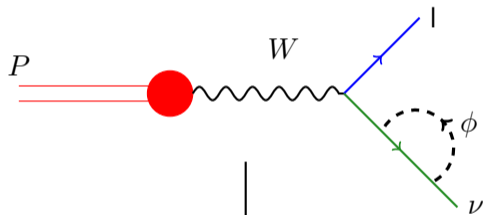
# It is not Divergent



# It is not Divergent!



+



$$\sum_s u^{(s)} \bar{u}^{(s)} \approx \not{p} \left[ 1 - \frac{g_l^2}{32\pi^2} \text{Log} \left( \frac{(m_P^2 - m_l^2)^2}{4m_\phi^2 m_P^2} \right) \right]$$

Using a  $\chi^2$ :

$$\chi^2 = \sum_{P,l} \frac{[\Gamma_{Pl}(1 + A_{Pl}g_l^2) - a_{Pl}]^2}{\sigma_{Pl}^2}$$

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$$\chi^2 = \sum_{P,l} \frac{[\Gamma_{Pl}(1 + A_{Pl}g_l^2) - a_{Pl}]^2}{\sigma_{Pl}^2}$$



Experimental Data 8 Decays:

$\pi, K, D^+, D_s, B$

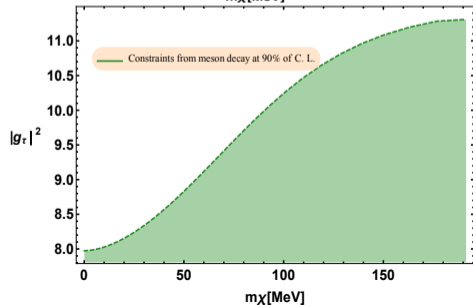
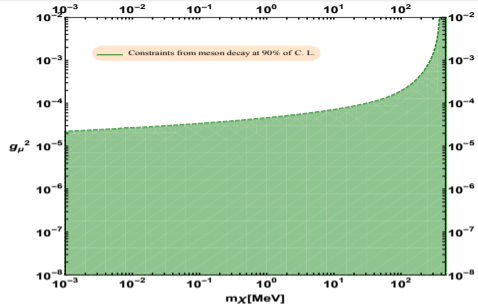
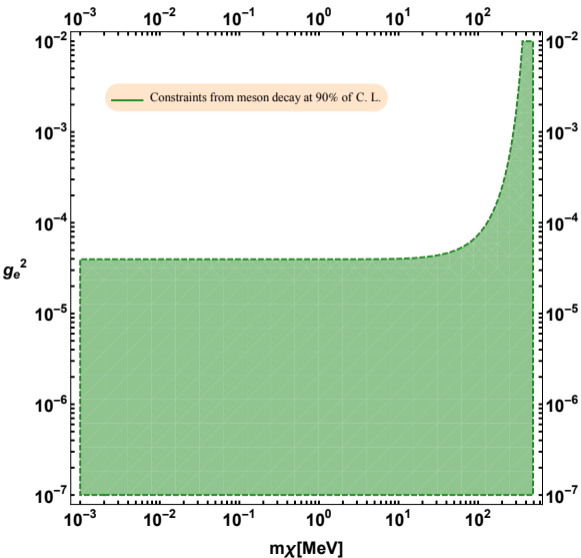
[5] PDG Chin. Phys. C 38, 090001 (2014).

[6] BESIII Collaboration, Phys. Rev. D 89, no. 5, 051104 (2014)

[7] BaBar Collaboration, arXiv:1003.3063 [hep-ex].

[8] Belle Collaboration, JHEP 1309, 139(2013).

# Bounds on $\pi, K, D^+, D_s, B$



## Continuous peaks = $l\nu\phi$ decay

Heavy Neutrino Search allows to use this calculations too!

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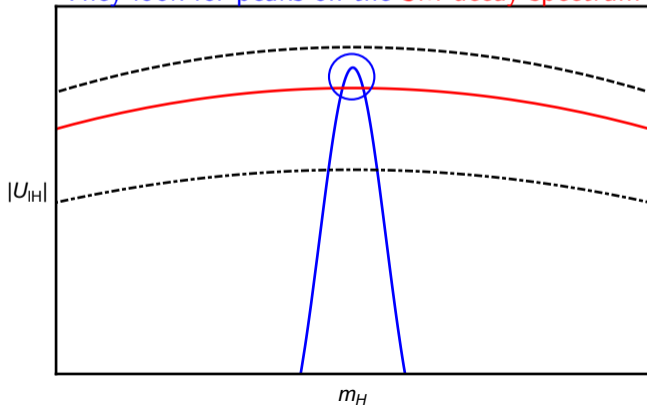


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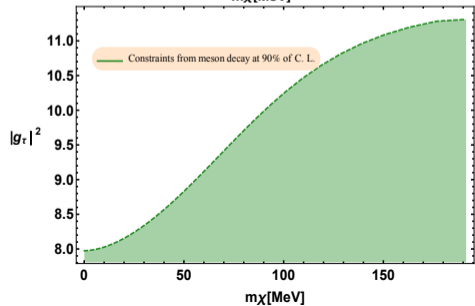
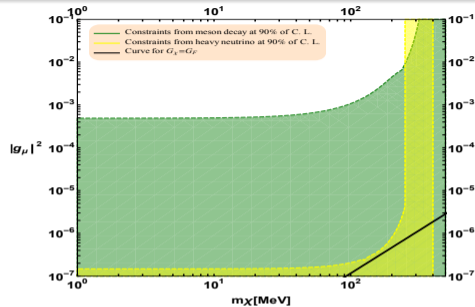
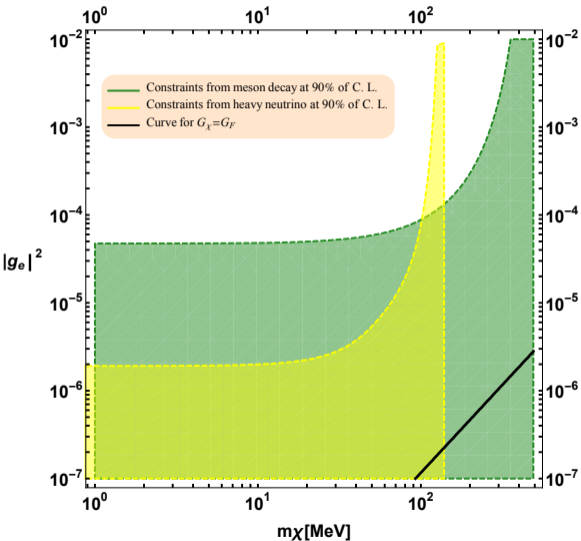
Heavy Neutrino Search allows to use this calculations too!

The three body decay mimic a continuous mass spectrum of heavy neutrino

They look for peaks on the SM decay spectrum



# Using Heavy Search from $\pi, K$ [8,9]



[8] D. I. Britton et al., Phys. Rev. D 46, 885 (1992).

[9] A. V. Artamonov et al. [E949 Collaboration], arXiv:1411.3963 [hep-ex].

# Heavy Search Uses Lepton Spectrum $\rightarrow$ much better

## Conclusions

We improved the bounds on the constants!

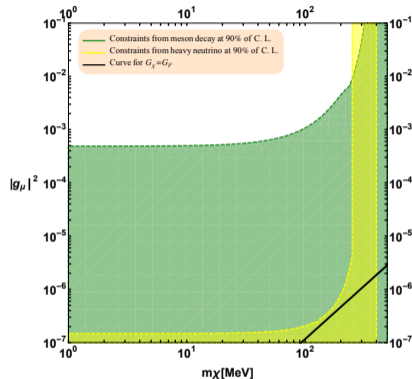
Our Results are valid up to  $m_\phi \sim 100$  MeV

Constants	Our Results	[10]	[11]
$ g_e ^2$	$< 1.9 \times 10^{-6}$	$< 4.4 \times 10^{-5}$	$< 1.6 \times 10^{-5}$
$ g_\mu ^2$	$< 1.9 \times 10^{-7}$	$< 3.6 \times 10^{-4}$	
$ g_\tau ^2$	$< 7.5$	$< 2.2 \times 10^{-1}$	

**Table:** Limits setting  $m_\phi = 0$ .

[10] A. P. Lessa and O. L. G. Peres, Phys. Rev. D 75, 094001 (2007).

[11] J. B. Albert et al. [EXO-200], Phys. Rev. D 90, no. 9, 092004 (2014).



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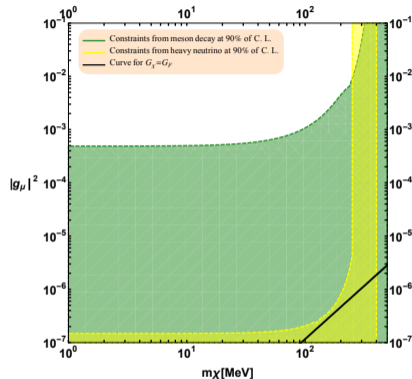
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**Agradecimentos:**

Fapesp

CNPQ

Unicamp (IFGW)

Advisor, Professors and friends

Backup Slides

# We need second order corrections..

Radiative Corrections<sup>2</sup>!

Tree Level

Lepton Radiative Corrections

$$\Gamma(P \rightarrow l\nu(\gamma)) = \Gamma^{(0)} \left(1 + 2\frac{\alpha_{el}}{\pi} \log\left(\frac{m_z}{m_\rho}\right)\right) \left[1 + \frac{\alpha_{el}}{\pi} F\left(\frac{m_l}{m_P}\right)\right] \times \left\{1 - \frac{\alpha_{el}}{\pi} G(m_\rho, m_P, m_l)\right\}$$



Low Distance Radiative Corrections

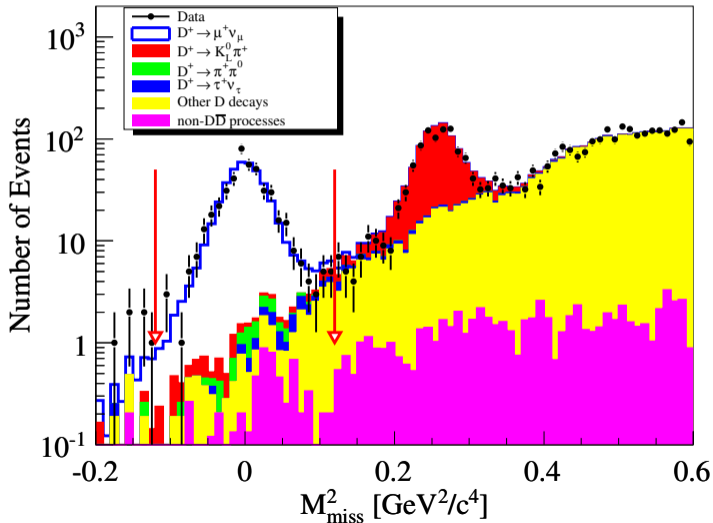


High Distance Radiative Corrections

<sup>2</sup>W. J. Marciano and A. Sirlin, Phys. Rev. Lett. **71**, 3629 (1993).

# Missing Energy Allows Probe New Physics

$$x \neq 0$$



# Bounds on $\pi, K, D^+, D_s, B$

