# The effects of SUSY seesaw on the dark matter and collider signals

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Based on the papers in collaboration with Keith Olive (Minnesota), Jing Shao (Syracuse), Liliana Vesasco-Sevilla (Cinvestav)

ulletModel:CMSSM(Constrained Minimal Supersymmetric Model)+  $N_3$ 

#### Dark Matter:

Features of the thermal relic abundance which don't show up in CMSSM.

- 1) Emergence of "Sneutrino Coannihilation Regions" (KK, K. Olive and L. Sevilla)
- 2) Complete disappearance of Focus Point Regions (KK and K. Olive)

#### Collider Signals:

Enhancement of the tau signals which don't show up in CMSSM.

3) 3 or more taus + jets (with negligible SM backgrounds) at 1/fb for small  $m_0$  and  $M_{1/2}$ .(KK and J. Shao)

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## Model

Toy model: Supersymmetric Seesaw

CMSSM+a right-handed neutrino  $N_3$  (GUT scale mass) (only two additional parameters:  $M_N(Q_{GUT}), m_v(Q_{Mz})$ )

Q: Does a heavy N affect the thermal dark matter abundance and/or collider signals? (in this talk, dark matter = neutralino LSP)

The coupling of *N* to the other fields:

$$y_N NLH_u$$

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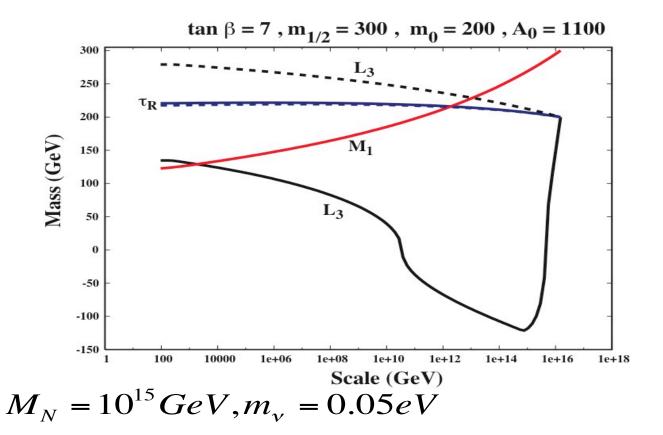
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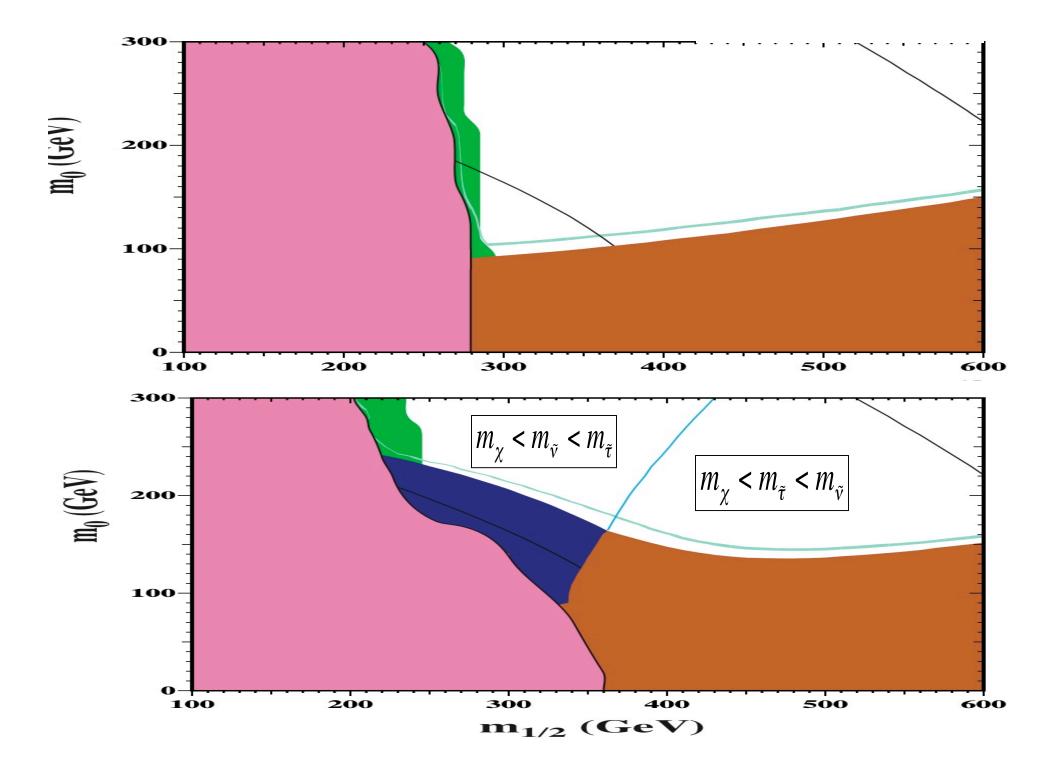
$$y_N NLH_u$$



Light left-handed sneutrino ⇒

Emergence of sneutrino coannihilation regions

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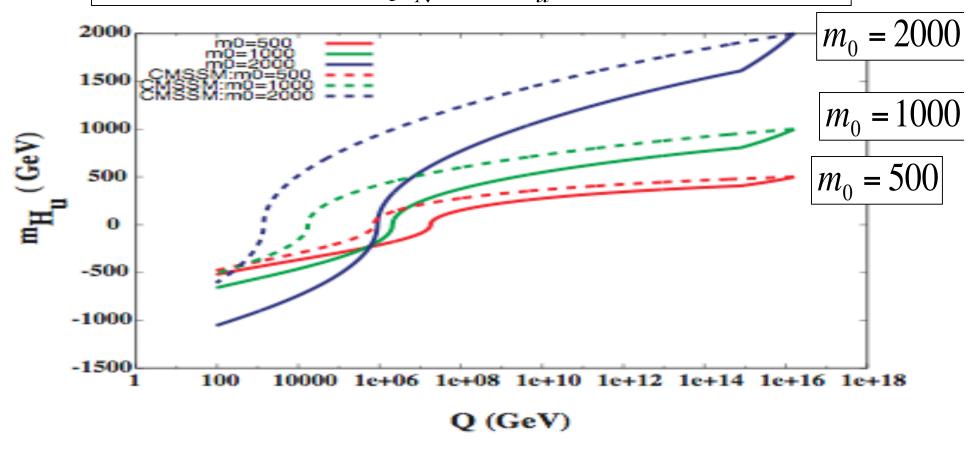
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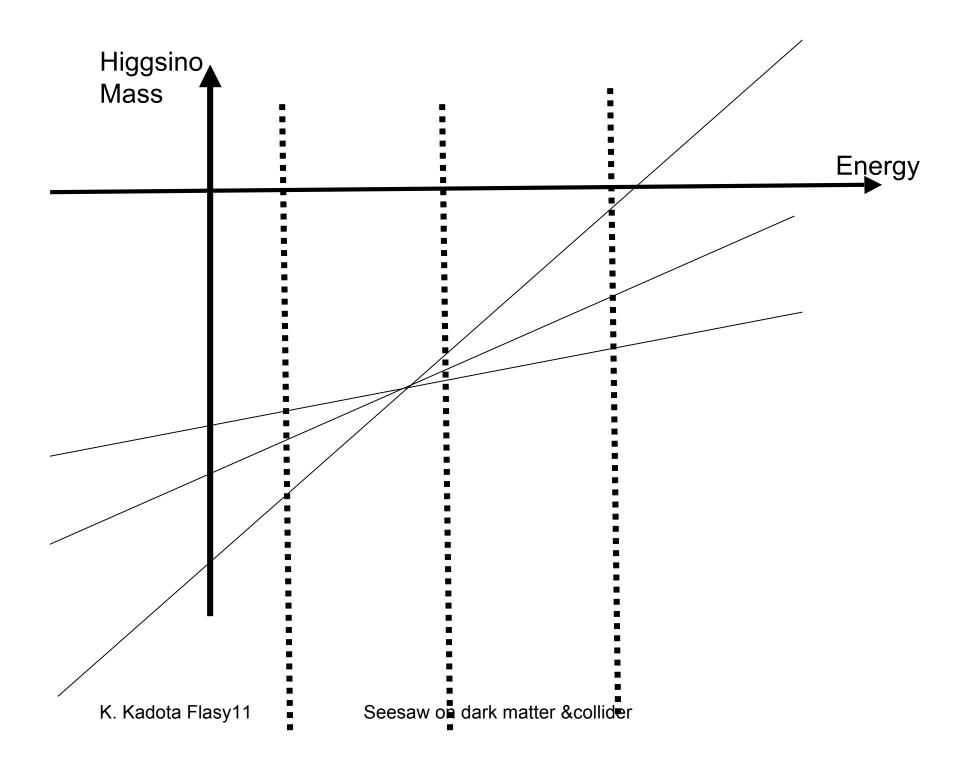
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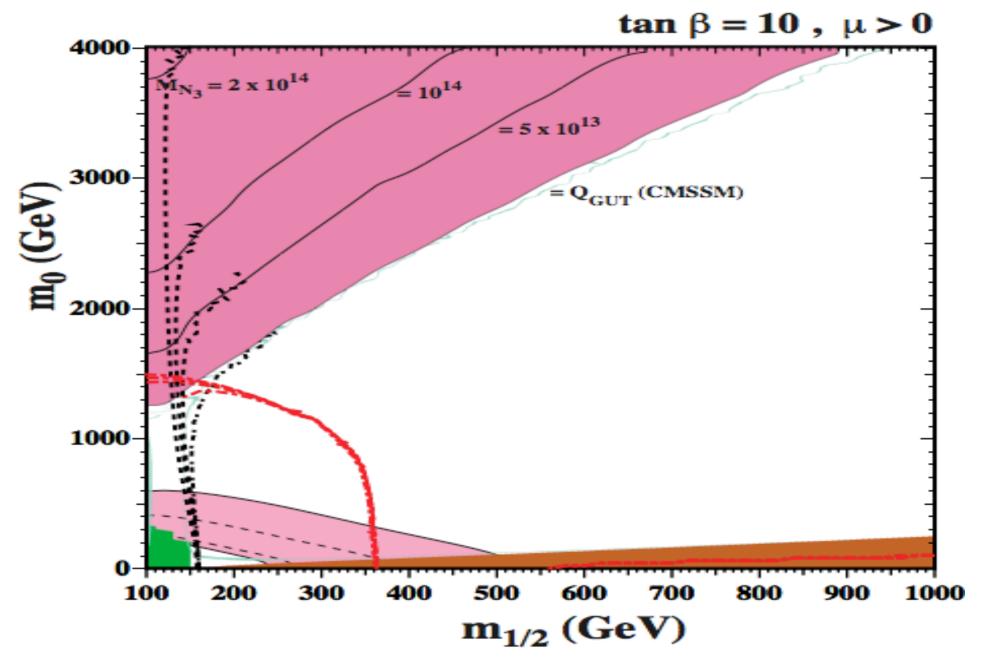
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## The coupling of N to the other fields: $y_N N L H_u$



Focus point scale can change dramatically ⇒ Disappearance of the focus point regions





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Seesaw on dark matter &collider

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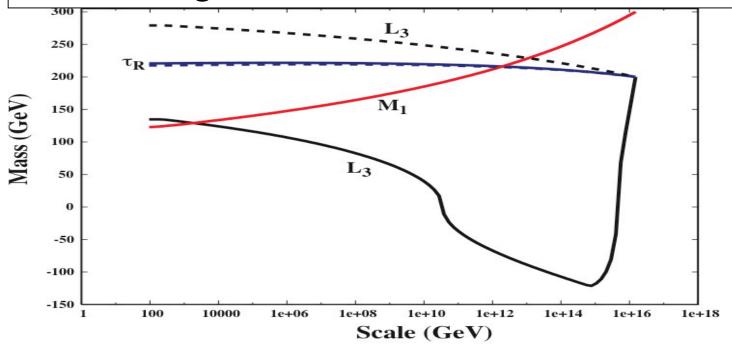
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#### Model: Light left-handed stau doublet



 $\tilde{\chi}_1^0$ : Bino - like.  $\tilde{\chi}_2^0$  &  $\tilde{\chi}_1^{\pm}$ : Wino - like.

CMSSM:  $\tilde{\tau}_1$  is RH

Drees&Nojiri '92, Baer et al '98, Barger&Kao '99, Lykken&Matchev '00, Hinchliffe&Paige '00, Wells '98, Arnowitt et al '06, Chattopadhyay et al '07, Katz&Tweedie '10 ...

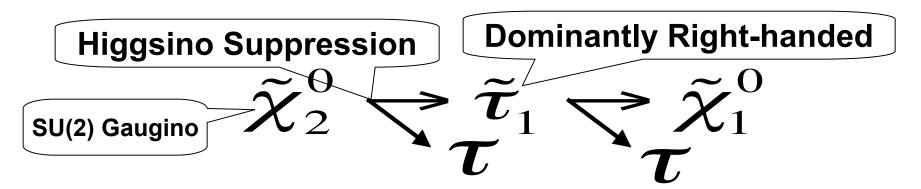
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KK&J. Shao '09

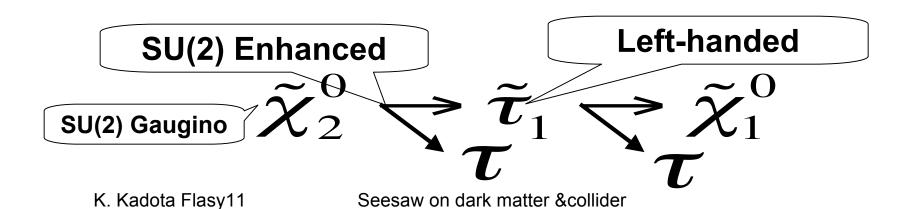
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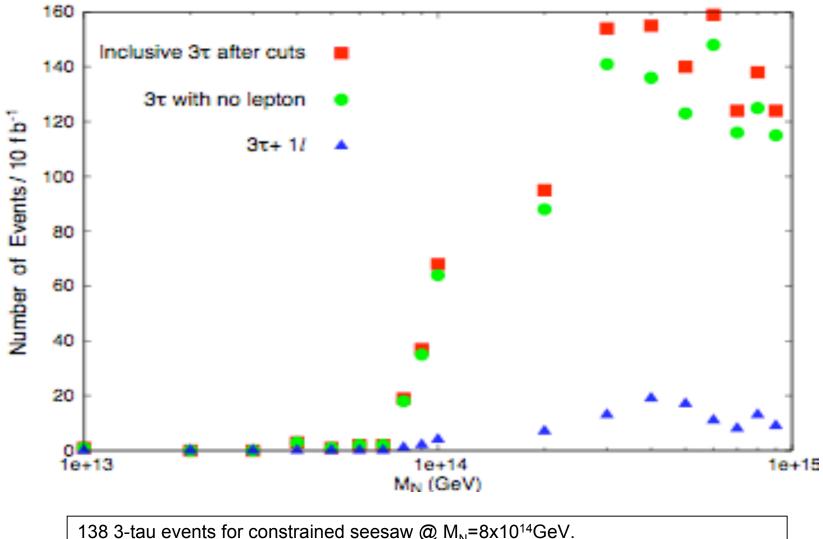
#### Enhanced tau events in light stau doublet scenario

Previous literature: Light <u>right-handed</u> stau
 Large tanβ or/and A<sub>0</sub>:Kinematically preferable small m<sub>0</sub>,m<sub>1/2</sub> tightly constrained by Br(b→sγ), Br(Bs→μ<sup>+</sup> μ<sup>-</sup>) etc.



This talk: Light <u>left-handed</u> stau
 Kinematically preferable small m<sub>0</sub>,m<sub>1/2</sub> (Sneutrino Co-annihilation region)





138 3-tau events for constrained seesaw @  $M_N$ =8x10<sup>14</sup>GeV. 1 3-tau event for CMSSM (tau efficiency factor  $\varepsilon_{\tau}$ =0.4.)

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<Work in progress: Inclusion of flavor mixings and CP phases>

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# Enhanced Tau Lepton Signatures in Constrained Supersymmetric Seesaw Scenarios

- Production and decay of weak gauginos could signal the deviations from SM e.g. Enhanced tau events
  - (could be related to finite neutrino masses)

$$W = W_{MSSM} + y_N NLH_u + \frac{1}{2} M_N NN$$
  

$$m_0, M_{1/2}, A_0, \tan \beta, sign(\mu)$$

 $M_N(Q_{GIIT}), m_{\nu}(Q_{M_7})$ 

- Simple cuts:
- > Four jets with pt>50GeV with the leading jets pt>100GeV
- MET>max(0.2Meff,100GeV)
- Hadronically decaying taus pt>20GeV, |η|<2.5.</p>

Tau efficiency factor  $\varepsilon_{\tau}$ =0.4.

Checked these strong cuts sufficiently reduce the backgrounds (ZZ, WZ, Z+Jets,WW,tt, QCD jets) to be negligible for 3 and 4 tau events for our study

• Jet rejection factor(function of  $\varepsilon_{\tau}$  and jet Et):

300 for 20GeV<Et<30GeV 500 for 30<Et<60GeV 1000 for 60<Et<100GeV 3000 for 100GeV<Et

- Tau reconstruction/identification
- E.g. to be tested by ~100/pb of LHC data for multi jet backgrounds

- Simple cuts:
- 1) Four jets with pt>50GeV with the leading jets pt>100GeV
- 2) MET>max(0.2Meff,100GeV)
- 3) Leptons (e, $\mu$ ) pt>20GeV and  $|\eta|$ <2.5
- 4) Hadronically decaying taus pt>20GeV,  $|\eta|$ <2.5. Tau efficiency factor  $\varepsilon_{\tau}$ =0.4.

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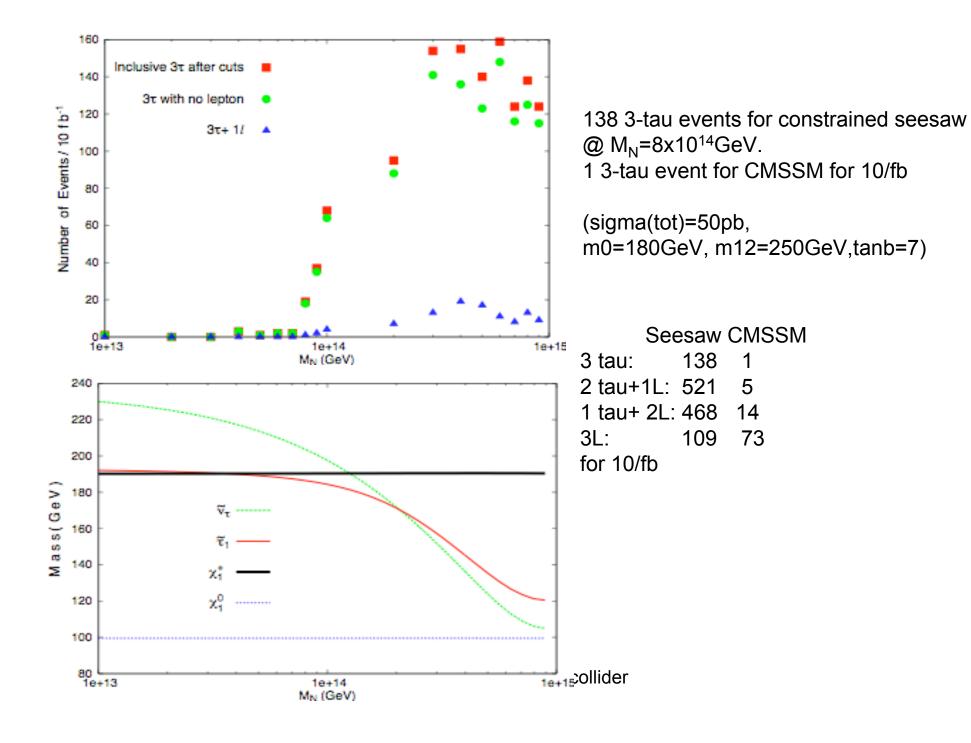
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## Tau Decay

- 35% leptonically with two neutrinos
- 65% hadronically with one (50%) or three (15%) charged particles

Low multiplicity and high collimation



## Model: Constrained Seesaw

#### CMSSM(Constrained Minimal Supersymmetric Model)+N

$$W = W_{MSSM} + y_N NLH_u + \frac{1}{2}M_N NN$$

$$m_0, M_{1/2}, A_0, \tan \beta, sign(\mu)$$

$$M_N(Q_{GUT}), m_v(Q_{Mz})$$

$$Q_{GUT} \sim 2 \times 10^{16} GeV, M_N \sim 10^{15} GeV$$

$$Q < M_N : L \ni -\kappa(LH_u)(LH_u) \Rightarrow m_v(Q_{EW}) = \kappa \langle H_u \rangle^2$$

Q: Does a heavy N affect collider signals?